

PRECESSION OF THE POLARIZATION OF PARTICLES MOVING IN A HOMOGENEOUS ELECTROMAGNETIC FIELD*

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The problem of the precession of the "spin" of a particle moving in a homogeneous electromagnetic field—a problem which has recently acquired considerable experimental interest—has already been investigated for spin $\frac{1}{2}$ particles in some particular cases.¹ In the literature the results were derived by explicit use of the Dirac equation, with the occasional inclusion of a Pauli term to account for an anomalous magnetic moment. On the other hand, following a remark of Bloch² in connection with the nonrelativistic case, the expectation value of the vector operator representing the "spin" will necessarily follow the same time dependence as one would obtain from a classical equation of motion. To solve the problem for arbitrary spin in the relativistic case, it will thus suffice to produce a consistent set of covariant classical equations of motion. Such equations have been indicated a long time ago by Frenkel³ and are discussed by Kramers.⁴ These authors use an antisymmetric tensor M as the relativistic generalization of the intrinsic angular momentum observed in the rest-frame of the particle. A formulation in terms of the (axial) four-vector s which describes the polarization in a covariant fashion⁵—though basically equivalent—is however much more convenient for our problem. We shall therefore derive first the equations of motion directly in terms of this four-vector s .

Let the spin of the particle be represented⁶ in the rest-frame (R) by $\vec{s}$. We assume (a) that there exists a four-vector s such that in (R) it coincides with $\vec{s}$:

$$s = (s^0, \vec{s}); \quad \text{in } (R), \quad s = (0, \vec{s}). \quad (1)$$

Denoting the four-velocity of the particle by $u = (u^0, \vec{u}) = \gamma(1, \vec{v})$ [where $\vec{v}$ is the ordinary velocity, and $\gamma(v) = (1 - v^2)^{-1/2}$], one has in every frame

$$s \cdot u = 0, \quad \text{i.e.,} \quad s^0 = \vec{s} \cdot \vec{v}. \quad (2)$$

We further assume (b) that $\vec{s}$ obeys in (R) the

customary equation of motion

$$d\vec{s}/d\tau = (ge/2m)(\vec{s} \times \vec{H}), \quad (R) \quad (3)$$

where $\vec{H}$, e , and m have their standard meanings, while the gyromagnetic ratio g is defined by this very equation. While s^0 vanishes by hypothesis in any instantaneous rest-frame, $ds^0/d\tau$ need not. In fact, (2) implies

$$ds^0/d\tau = \vec{s} \cdot (d\vec{v}/d\tau), \quad (R) \quad (4)$$

for such frames. In general, $du/d\tau = f/m$ (where f = four-force), while in a homogenous external electromagnetic field specified by $F = -(\vec{E}, \vec{H})$

$$du/d\tau = (e/m)F \cdot u. \quad (5)$$

The immediate generalization of Eqs. (3) and (4) to arbitrary frames is

$$ds/d\tau = (ge/2m)[F \cdot s + (s \cdot F \cdot u)u] - [(du/d\tau) \cdot s]u, \quad (6)$$

as can be checked by reducing to the rest-frame. With (5), one has for homogeneous fields

$$ds/d\tau = (e/m)[(g/2)F \cdot s + (g/2 - 1)(s \cdot F \cdot u)u]. \quad (7)$$

(5) and (7) constitute, for any value of g and arbitrary spin s , a consistent set of equations of motion; they imply that $s \cdot s$ and $s \cdot u$ are constant, so that condition (2) is maintained.⁷ For experiments of current interest, the main use of (7) is in the computation of the rate Ω at which longitudinal polarization is transformed into a transverse one (and vice versa). For this, we express s in the laboratory frame (L) in terms of two unit polarization four-vectors, e_L and e_T :

$$s/s = e_L \cos \phi + e_T \sin \phi,$$

where

$$s = (-s \cdot s)^{1/2},$$

$$e_L = \gamma(v, \vec{v}/v) \equiv \gamma(v, \hat{v}), \quad e_T = (0, \hat{n}),$$

$$\hat{n} \cdot \hat{n} = 1, \quad \hat{n} \cdot \hat{v} = 0. \quad (L) \quad (8)$$

Clearly, $\Omega = d\phi/dt = d\phi/\gamma d\tau$. Introducing (8) into (7), and expressing all quantities as ordinary vectors, we find

$$\Omega = (e/m)\{(\vec{E} \cdot \hat{n}/v)[(g/2 - 1) - g/2\gamma^2] + (\hat{v} \cdot \vec{H} \times \hat{n})(g/2 - 1)\}. \quad (9)$$

The relevant "anomaly" of spin- $\frac{1}{2}$ particles, $(g/2 - 1)$, is clearly exhibited in (9) although our derivation was classical throughout.

We now specialize (9) to some cases of practical interest (the references are to experiments):

(A) $\vec{E} \times \vec{v} = \vec{H} \times \vec{v} = 0$; $\Omega = 0$. The character of the polarization does not change, but the transverse polarization precesses around $\vec{v}$ in longitudinal fields with an angular frequency $\omega = (ge/2m\gamma)H = (g/2)\omega_L$, as follows readily from (8).

(B)⁸ $\vec{H} \cdot \hat{n} \times \hat{v} = H$, $\vec{E} = 0$; $\Omega = \omega_L(g/2 - 1)\gamma$, where ω_L is the Larmor frequency defined in (A) above.

(C)⁹ $\vec{E} \cdot \hat{n} = E$, $\vec{H} = 0$; $\Omega = \omega_p[-g/2\gamma + (g/2 - 1)\gamma]$, where $\omega_p = eE/m\gamma v$ is the angular frequency of the particle's motion in the laboratory.

(D)¹⁰ $\vec{E} \cdot \vec{H} = 0$, rectilinear motion: $\vec{E} = -\vec{v} \times \vec{H}$; $\vec{H} \cdot \hat{n} \times \hat{v} = H$, $\Omega = \omega_L(g/2\gamma)$.

(E)¹¹ $\vec{E} \cdot \vec{H} = 0$, $\vec{H} \cdot \hat{n} \times \hat{v} = H$, $\vec{E} \cdot \hat{n} = -Ev_x/v$, trochoidal motion: $E/H < 1$; $\Omega = (e/m)[(g/2 - 1)(H - Ev_x) + Ev_x/(\gamma^2 - 1)]$,

$$(\Delta\phi/2\pi) \text{ per loop} = \gamma(E/H)\gamma(v)[1 - (E/H)v_x](g/2 - 1) \\ = \gamma(v')(g/2 - 1),$$

where v' is velocity in a frame where $E' = 0$.

The generalization of (6) to cover particles having an intrinsic electric dipole moment $\vec{\epsilon} = (ge/2m)\vec{s}$ may be of interest. In the (R) frame, the effect of $\vec{\epsilon}$ is taken into account by adding $\vec{\epsilon} \times \vec{E}$ to the right-hand side of (3), while leaving (4) unchanged. Thus the required change in the right-hand side of (6) is the addition of a term $-(ge/2m)[(F^* \cdot s) + (s \cdot F^* \cdot u)u]$, denoting by F^* the dual of F , i.e., $F^* = -(\vec{H}, -\vec{E})$. For the experiment (B) above, one obtains then $|\Omega| = \omega_L\gamma \times [(g/2 - 1)^2 + (g'v/2)^2]^{1/2}$.

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¹H. A. Tolhoek and S. R. de Groot, *Physica* **17**, 17 (1951); H. Mendlowitz and K. M. Case, *Phys. Rev.* **97**, 33 (1955); L. M. Carrassi, *Nuovo cimento* **7**, 524 (1958).

²F. Bloch, *Phys. Rev.* **70**, 460 (1946).

³J. Frenkel, *Z. Physik* **37**, 243 (1926).

⁴H. A. Kramers, *Quantum Mechanics* (North Holland Publishing Company, Amsterdam, 1957), p. 226 et seq. Kramers' Eq. (4) (p. 229) does not correspond to our Eq. (6). His conclusion that already classically $g=2$ is implied for an electron, based on the relativistic equation he uses, stems from the fact that that equation corresponds in the rest-frame to $ds^0/d\tau = (ge/2m)\vec{s} \cdot \vec{E}$. Comparing with our (4), with $d\vec{v}/d\tau = (e/m)\vec{E}$, one sees that the "derived" result is, in fact, built into the theory from the start. A more general equation is mentioned by Kramers on p. 231, in fine print, and attributed to Frenkel. The inconsistencies arising in Kramers' discussion of what he calls "spin-orbit" forces [i.e., of the form $(\nabla\vec{H}) \cdot \vec{u}$ in the rest-frame] are connected with the fact that neither of his equations of motion applies when the field is inhomogeneous. In that case, $du/d\tau$ is not given by (5) alone, but has to include an additional term (the covariant analog of the gradient force just mentioned) before being introduced into (6).

⁵The covariant polarization four-vector s is essentially the expectation value of the operator w used by Bargmann and Wigner [Proc. Natl. Acad. Sci. U. S. **34**, 211 (1948)] to characterize representations of the inhomogeneous Lorentz group: $\langle w \cdot w \rangle = -s(s+1)m^2$, $s = \text{spin}$. s can be expressed in terms of the skew tensor M of Frenkel (which satisfies $M \cdot u = 0$), and vice versa: $s = M^* \cdot u$, $M^* = s \times u$, i.e., $M^*ik = s^iuk - s^kui$. For the quantum-mechanical applications of s see, e.g., C. Bouchiat and L. Michel, *Phys. Rev.* **106**, 170 (1955).

⁶Our notation is: $c=1$, $\hbar=1$ throughout; coordinate four-vector of components $x^0=t, x^1, x^2, x^3$: $x=(x^0, \vec{x})$, $\vec{x}=\{x^\alpha\}$ ($\alpha=1, 2, 3$); metric of signature $(+---)$; τ =proper time; a dot between symbols, contraction of neighboring indices with the metric tensor, e.g., $x \cdot x = (x^0)^2 - \vec{x}^2$; skew tensor of components T^{ik} indicated as $T=(\vec{T}^t, \vec{T}^s)$, $\vec{T}^t=\{T^{0\alpha}\}$, $\vec{T}^s=\{T^{\beta\gamma}\}$; $\alpha, \beta, \gamma=1, 2, 3$; its dual by $T^*=(\vec{T}_s, -\vec{T}_t)$.

⁷Equations (5) and (7) can be integrated explicitly by reference to four orthonormal four-vectors $e^{(i)}$ such that each of them obeys (5), and $e^{(0)}=u$.

⁸Crane, Pidd, and Louisell, *Bull. Am. Phys. Soc. Ser. II*, **3**, 369 (1958).

⁹H. Frauenfelder et al., *Phys. Rev.* **106**, 386 (1957).

¹⁰P. E. Cavanagh et al., *Phil. Mag.* **2**, 1105 (1957).

¹¹P. S. Farago, *Proc. Phys. Soc. (London)* **72**, 891 (1958).